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Elektromos sugárzás energiája

$$E = h\nu = hc\frac{1}{\lambda}. \quad (1.1)$$

Planck's constant $h = 6.626 \times 10^{-34}$ Js; the speed of light $c \approx 3 \times 10^8$ m/s.

$$hc \approx 19.86 \times 10^{-26} \text{ Jm}$$

$$1\text{eV} \approx 1.6 \times 10^{-19} \text{ J}$$

$$\mathbf{E[eV]} \approx \mathbf{1.24 \times 10^{-6} \cdot \frac{1}{\lambda[m]} eVm.}$$

Example: a lower energy gamma photon's wavelength is $\approx 10^{-12} \text{ m}$, thus:

$$E_\gamma[\text{eV}] = 1.24 \times 10^{-6} \cdot \frac{1}{10^{-12}} = 1.24 \text{ MeV.} \quad (1.2)$$

The energies of visible photons are in the range of: $1.6 \text{ eV} \rightarrow 3.4 \text{ eV}$, thus approximately $E_\gamma = 10^6 \times E_{vis}$.

Thermal energy, at room temperature: $T=20 \text{ }^\circ\text{C}$: $1/40 \text{ ev} = 25 \text{ meV}$.

Energy temperature correspondence:

$$E = k_b \cdot T, \quad (1.3)$$

where k_b is the Boltzmann-constant $k_b = 1.38 \times 10^{-23} \text{ J/K}$.

The temperature can be calculated as:

$$T[K] = \frac{E[\text{eV}]}{k_b} \quad (1.4)$$

For visible light ($E_{vis} \approx 1.6 \text{ eV}$) $T \approx 10^4 \text{ K}$.

1 eV $\rightarrow 10^4 \text{ K}$.